

2004  
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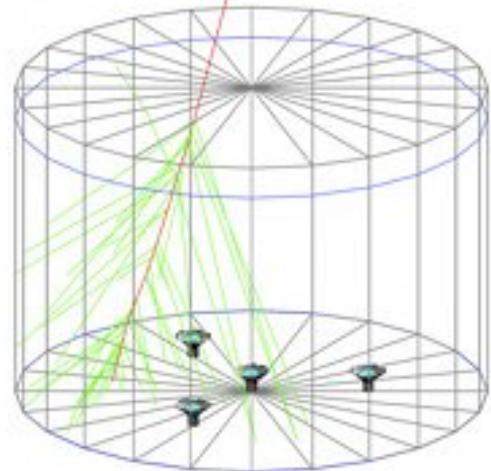


**GRAN  
TELESCOPIO**

# HAWC



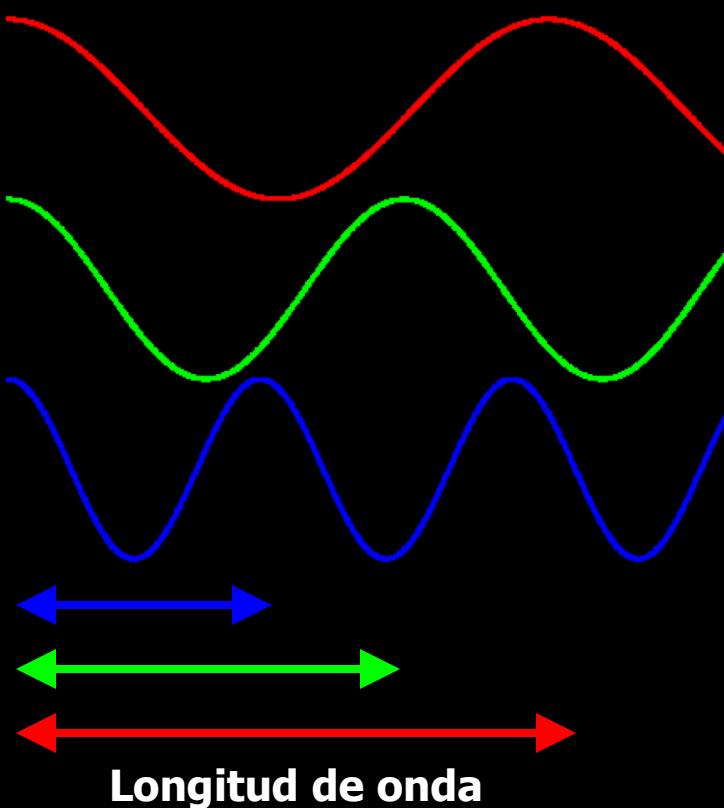
# HAWC



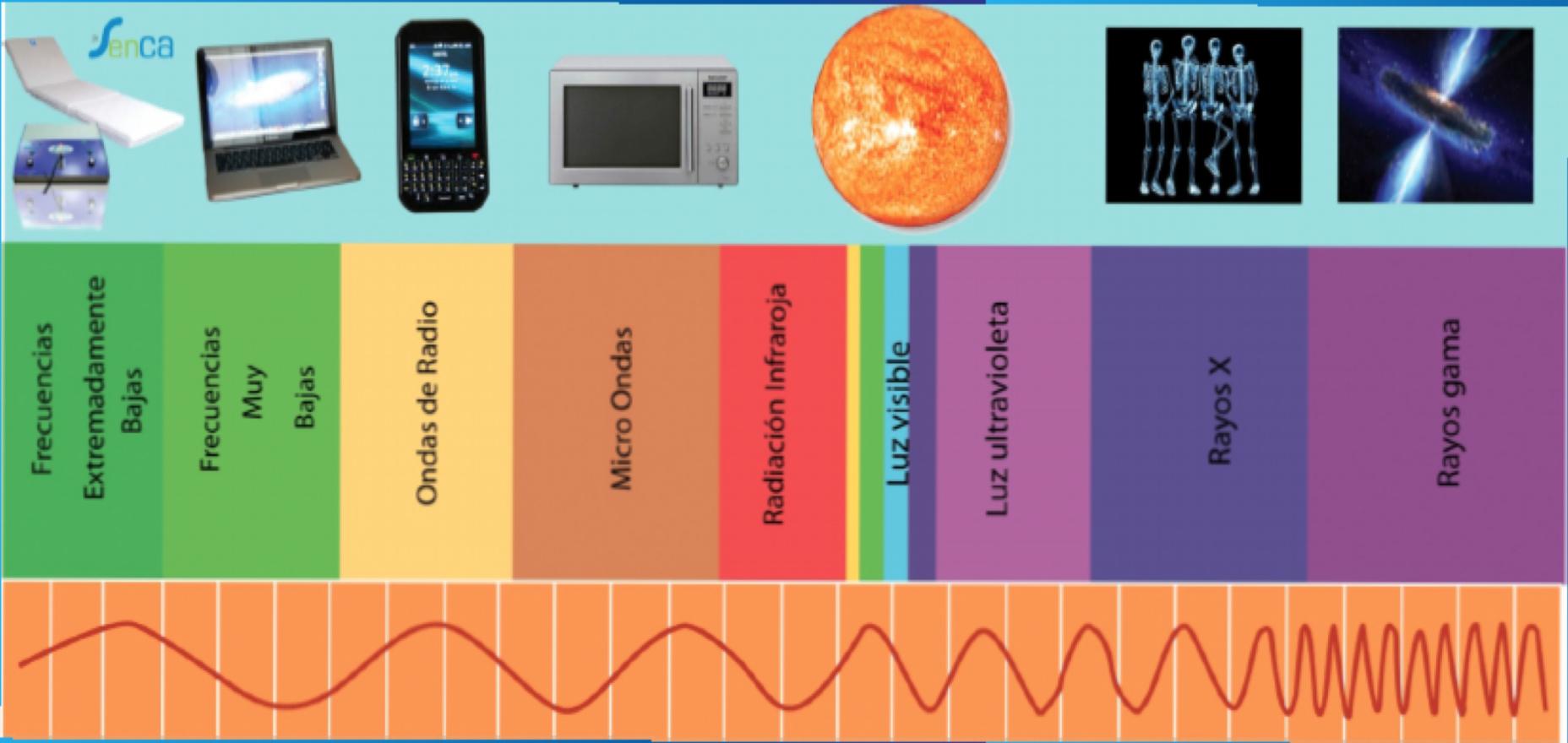
# Multiwavelength Astronomy

# Radiación Electromagnética

- Se describe la radiación EM por:
  - Frecuencia
  - Longitud de onda
- Luz:
  - 300,000 km/s



# Espectro Electromagnético



¿Como vemos la luz que nos llega del Universo?



Arecibo



GTM



Spitzer



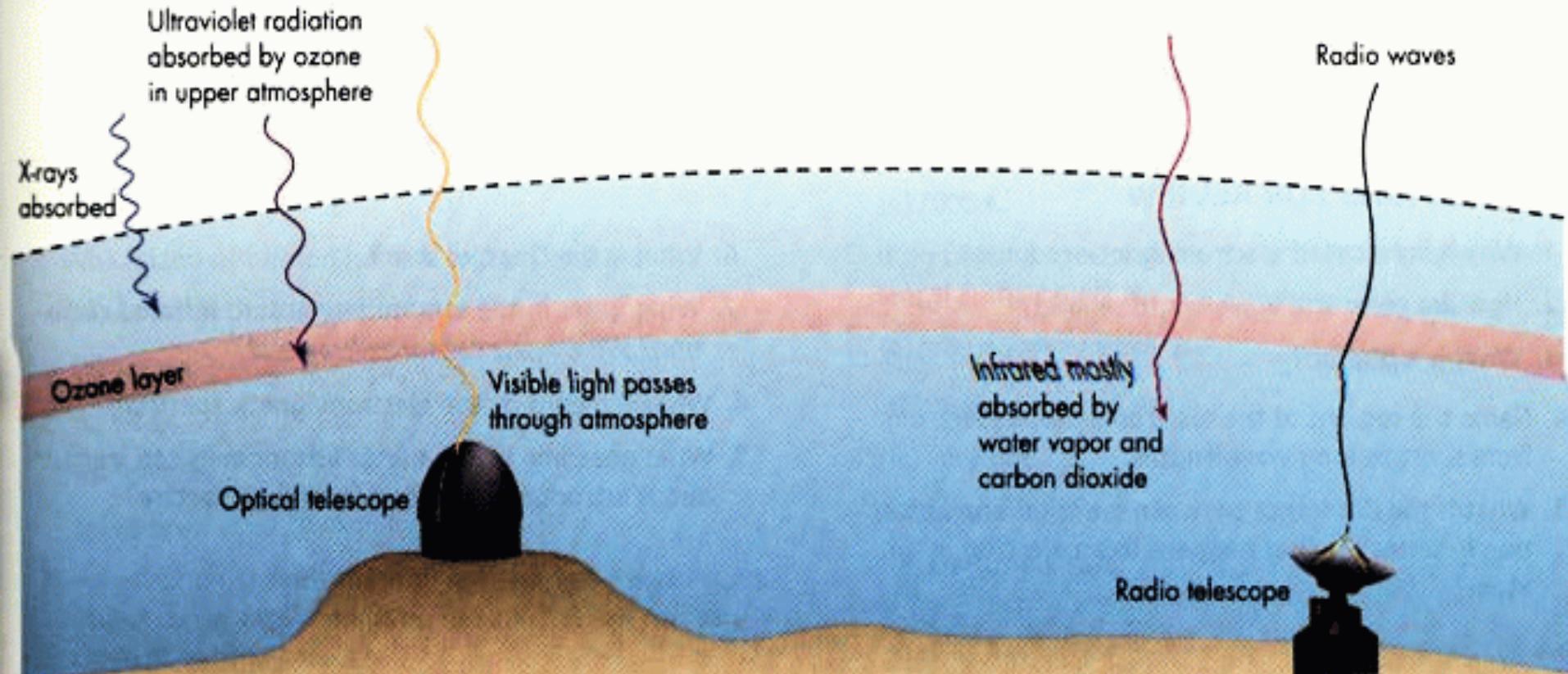
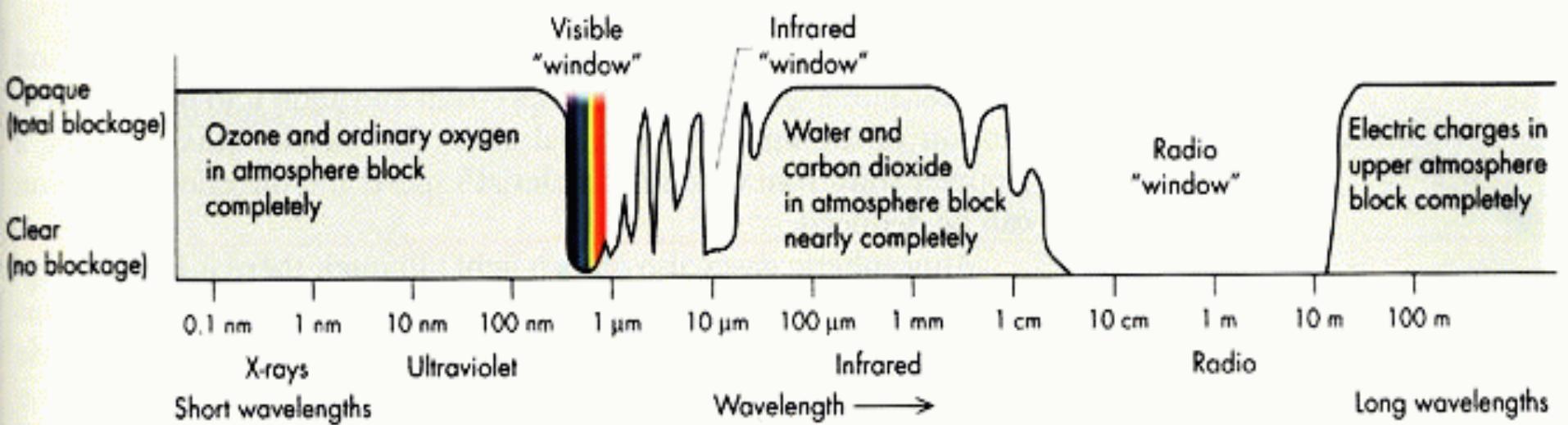
Hubble



Chandra



HAWC

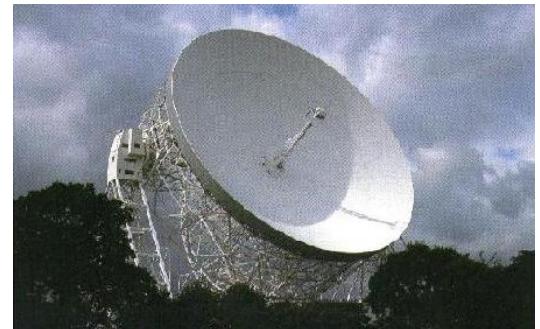
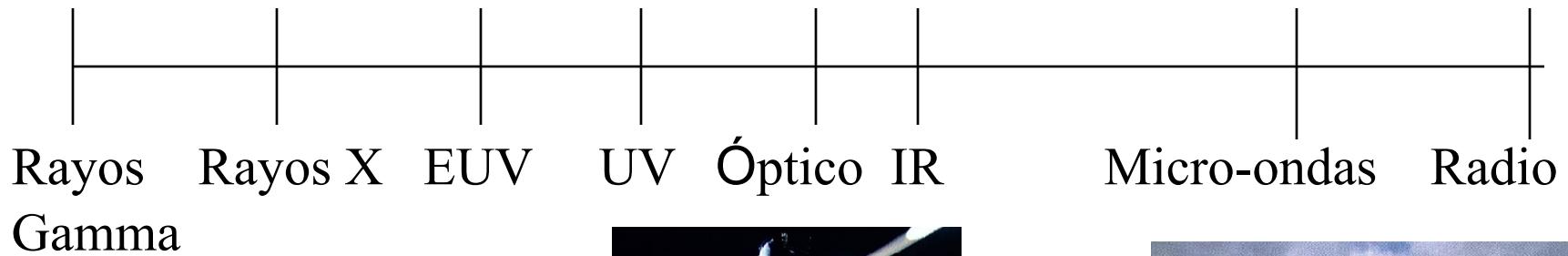


# *El Espectro Electromagnético*

100MeV 100keV 0.1keV    1000A 5000A 1micra

1cm

100m



# Multi-wavelength Astronomy

Banda	Longitud de Onda /Energía	Temperatura	Qué se puede estudiar
Rayos Gamma	100keV-100MeV	$>10^8$ K	Discos de acreción, gamma-ray bursts
Rayos X	<1-100keV	$10^6$ - $10^8$ K	Gas caliente en cúmulos de galaxias, coronas estelares, Discos de acreción,
Ultravioleta (UV)	900-3000Å	$10^4$ - $10^6$ K	Estrellas calientes enanas blancas, gas interestelar,
Óptico	3000-10,000Å	$10^3$ - $10^4$ K	Estrellas tipo solar
Infrarrojo	1-100 micras	$10$ - $10^3$ K	Polvo, planetas, enanas cafés
Microondas	1cm	<10K	Radiación de fondo del Universo.
Radio	>1m	<10K	Radiación de electrones moviéndose en un campo magnético: pulsars, AGNs

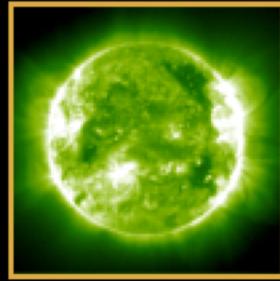
# Herschel



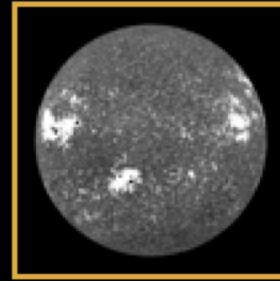
# El Sol



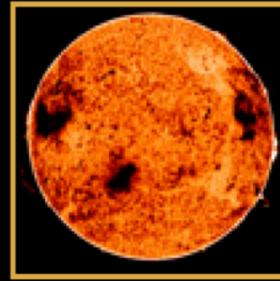
X-Ray: Yohkoh



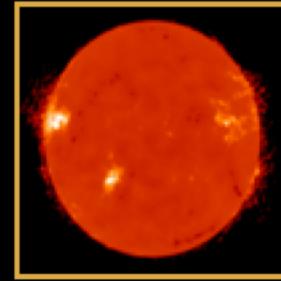
Ultraviolet: SOHO-EIT



Visible: BBSO

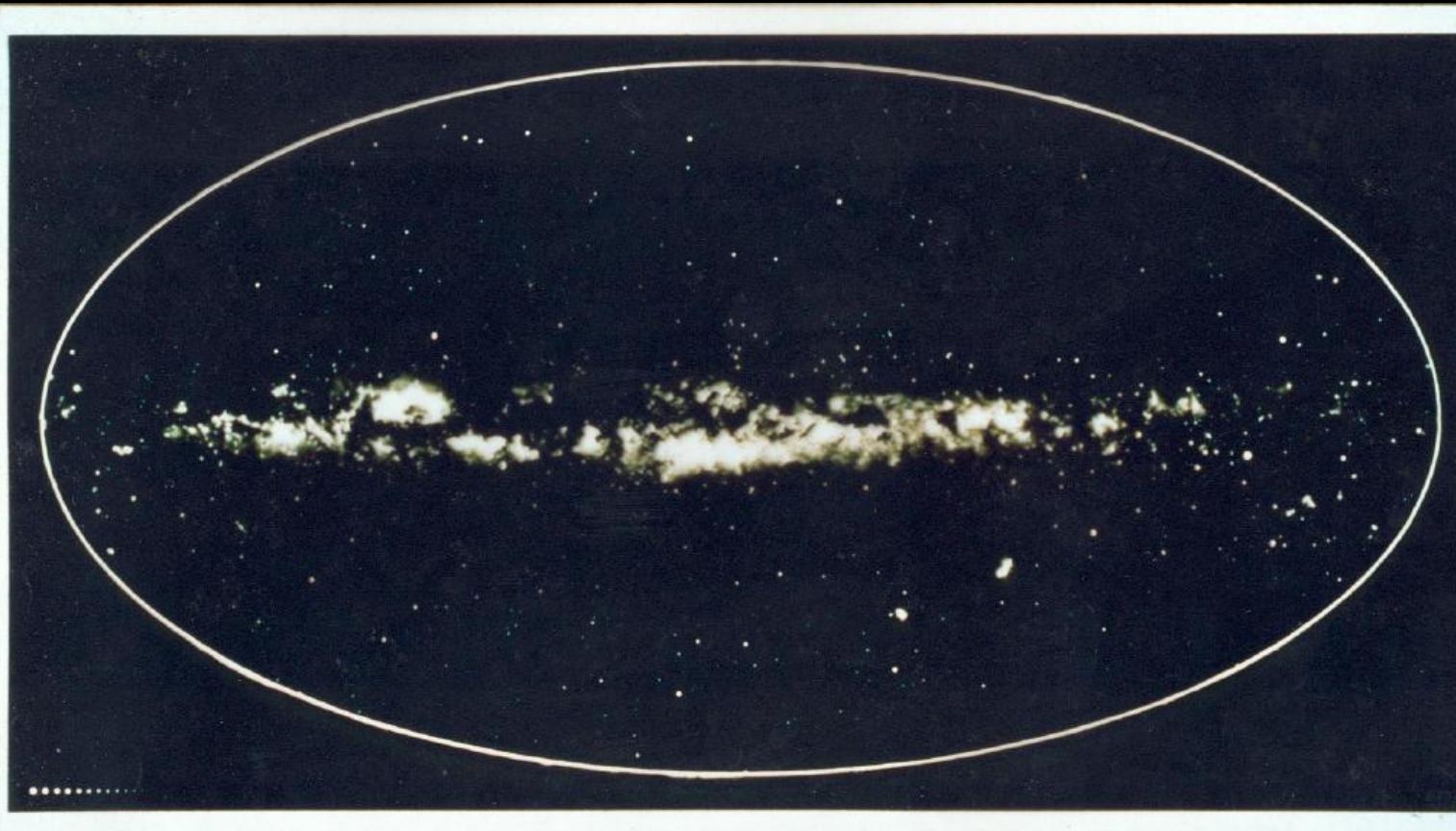


Infrared: NSO



Radio: Nobeyama

# La Vía Láctea: Óptico



estrellas,  
polvo

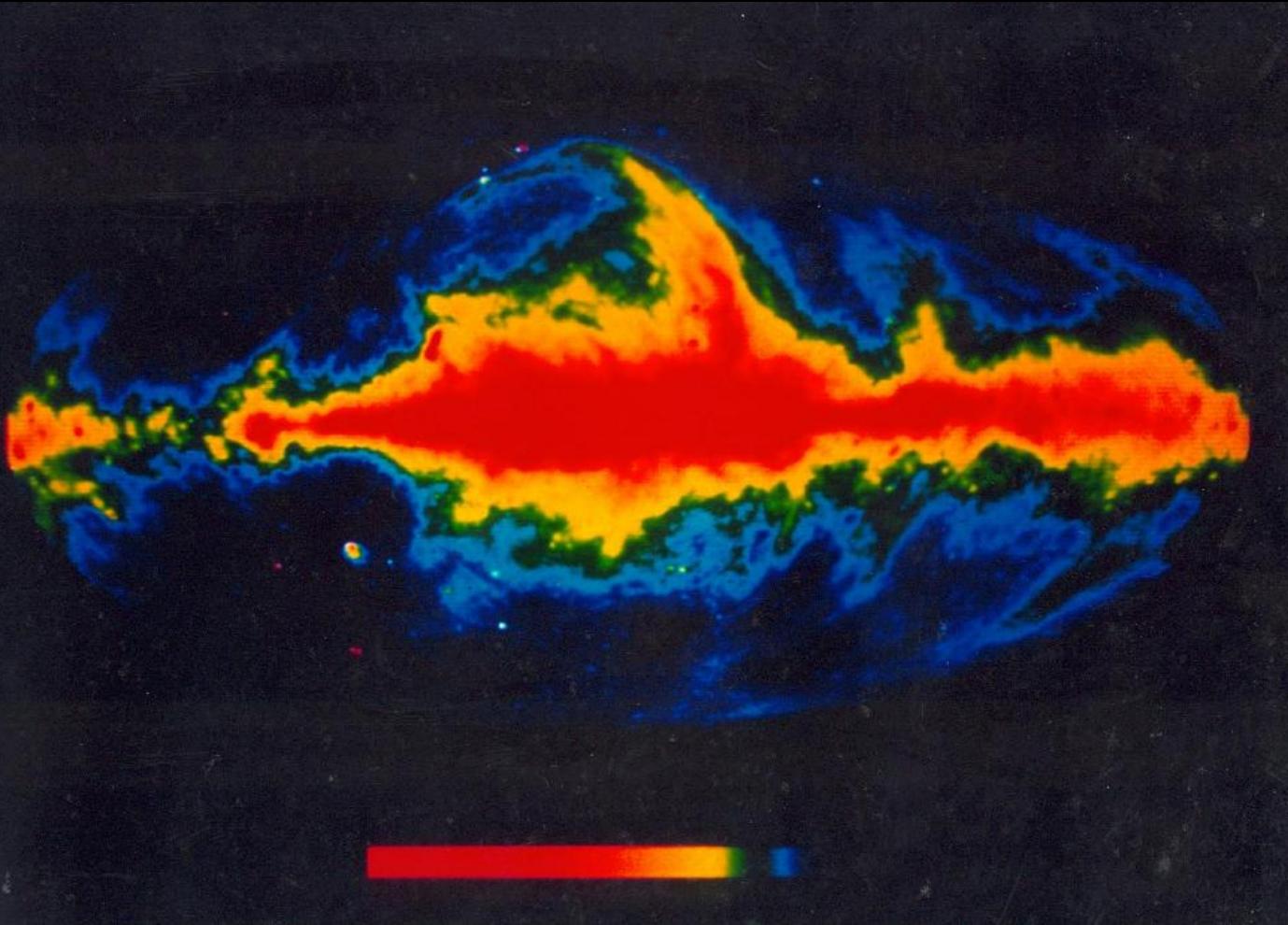
# La Vía Láctea: IR



*Near-infrared Image of the Milky Way obtained by the DIRBE on NASA's satellite COBE. NASA/Goddard Space Flight Center, April 17, 1990.*

A través  
del  
polvo

# La Vía Láctea: rayos X



Gas  
caliente

**Radio**

**Microondas**

**Infrarrojo lejano**

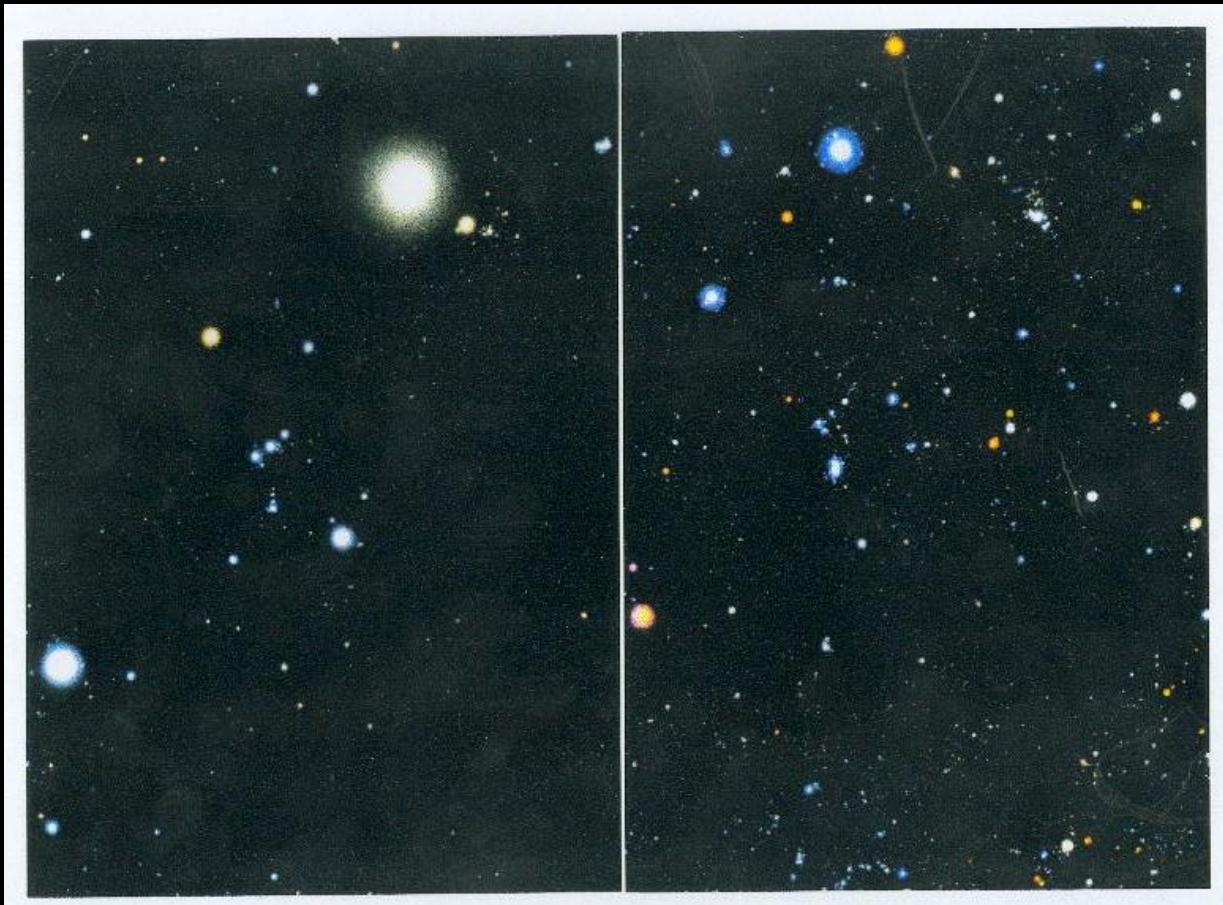
**Infrarrojo cercano**

**Óptico**

**Ultravioleta**

**Rayos X**

# Orión en el óptico y rayos X



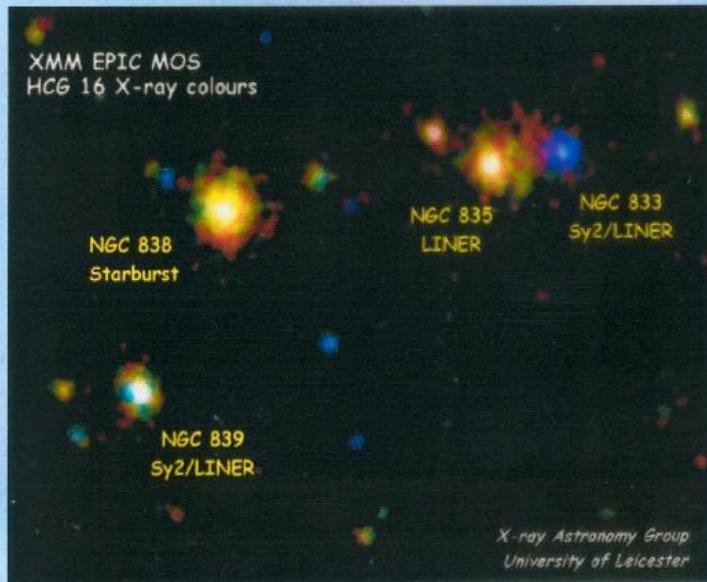
óptico

rayos X

# Estudio Multi-frecuencia

Hickson 16 – a compact group of AGN/Starburst galaxies

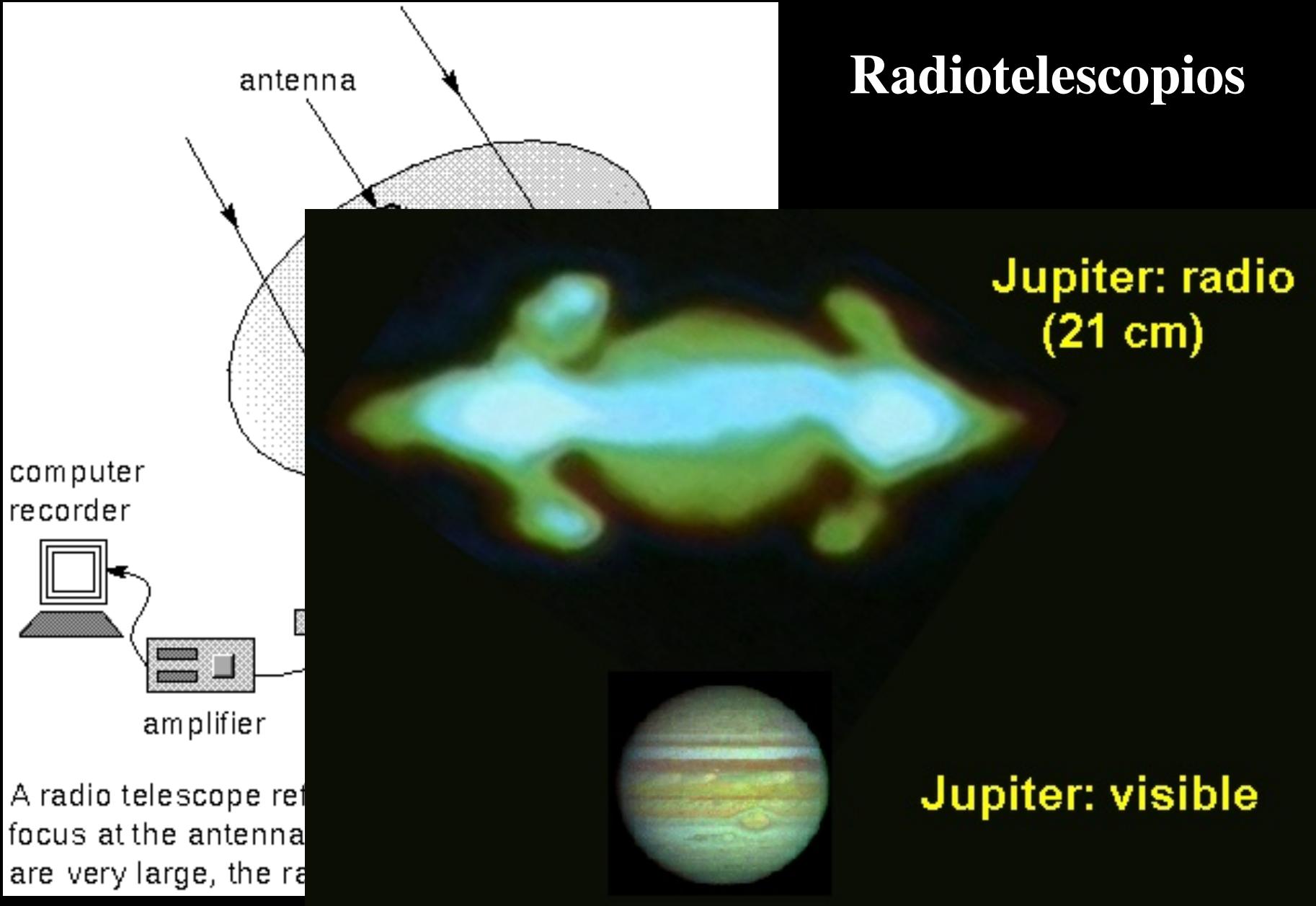
XMM X-ray



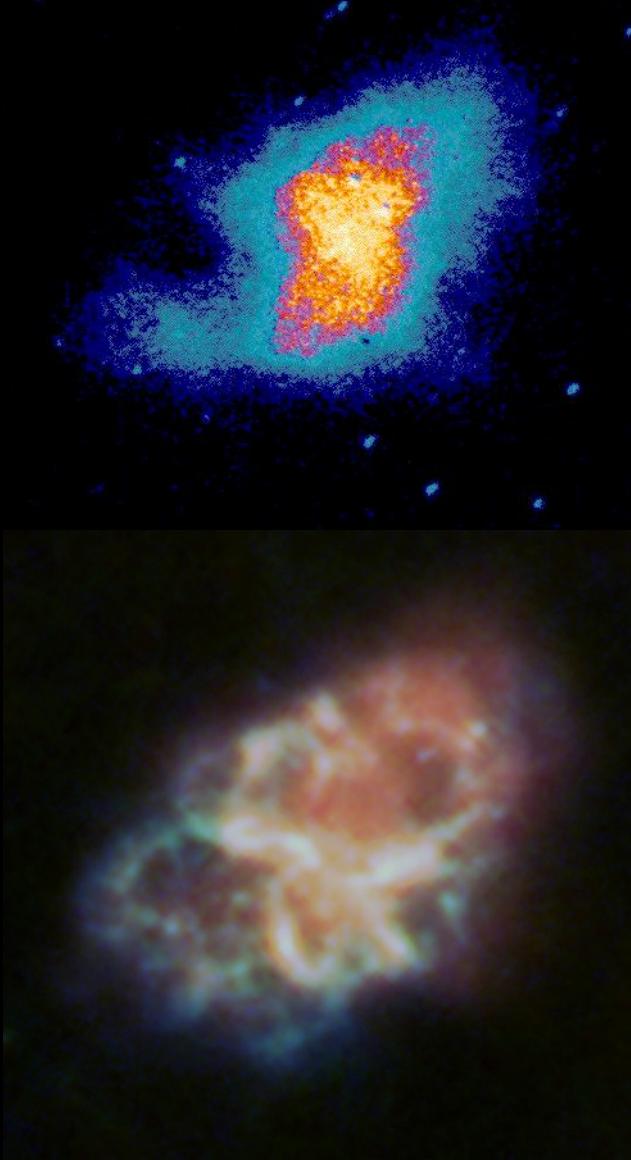
Optical



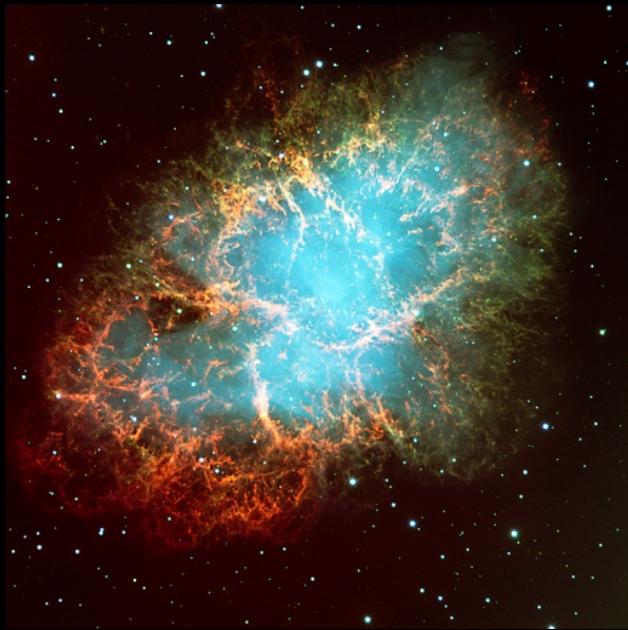
# Radiotelescopios



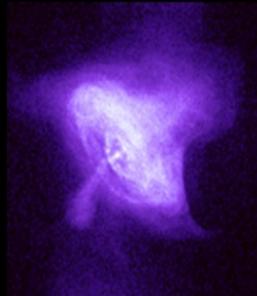
# Nebulosa del Canareio



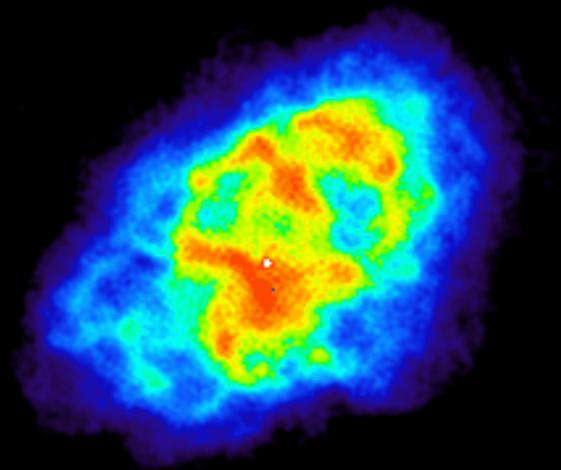
F7 (Herschel)



Crab (VLT)



X4 (Chandra)

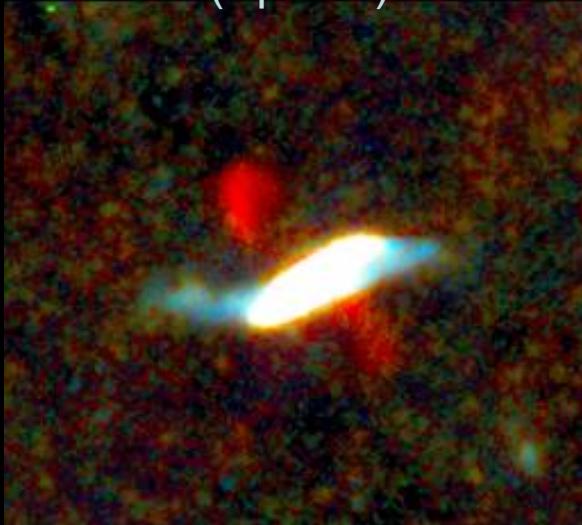


R9 (NRAO)

# Centaurus A



M7 (Spitzer)



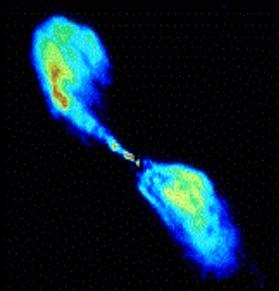
F8 (Herschel)



Centaurus A (NOAO)



X3 (Chandra)



R7 (NRAO)

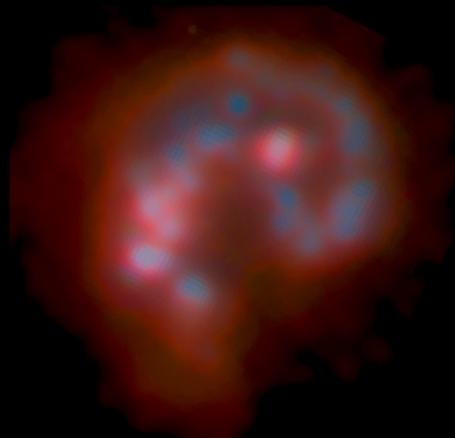
# Antennae



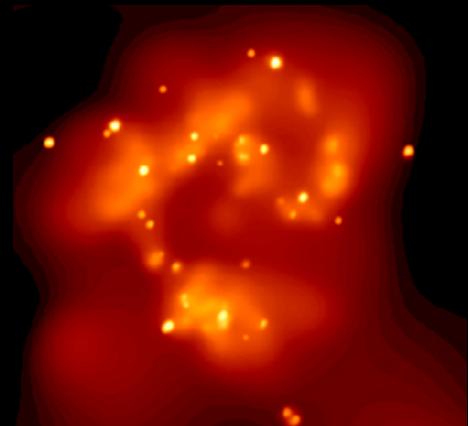
**M6** (Spitzer)



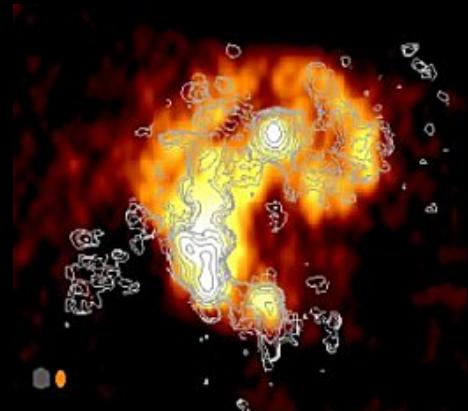
Antennae (Hubble)



**F1** (Herschel)

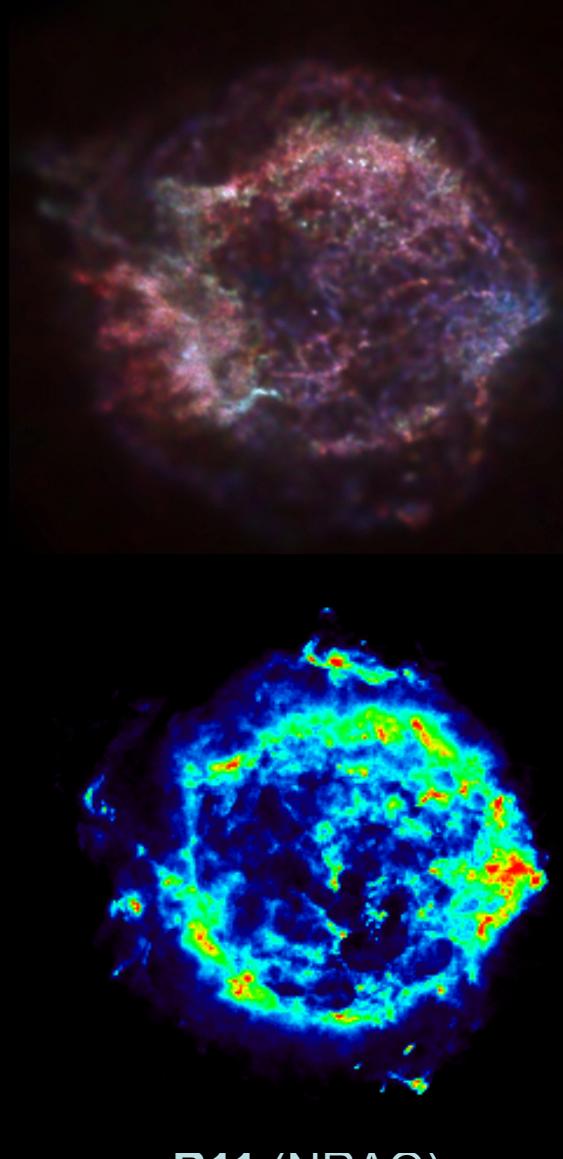
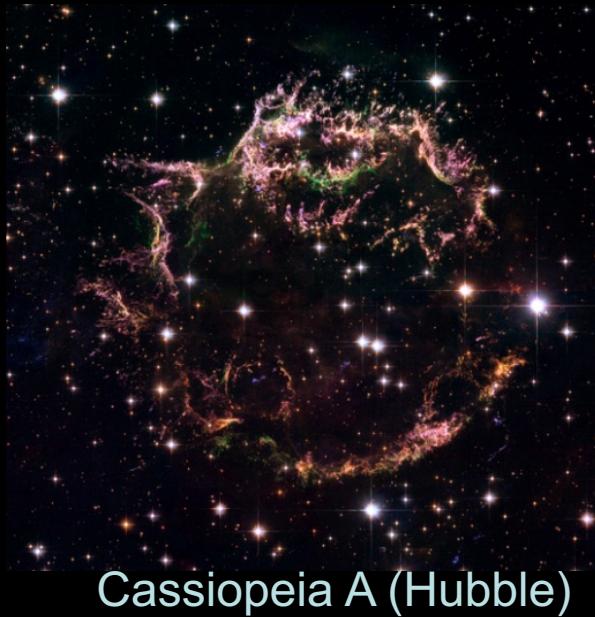


**X5** (Chandra)

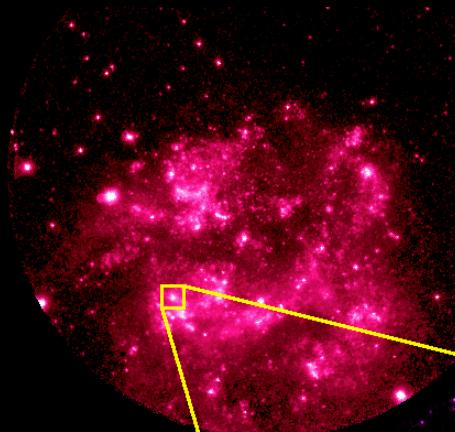


**R8** (VLA)

# Cassiopeia A



# Gran Nube de Magallanes



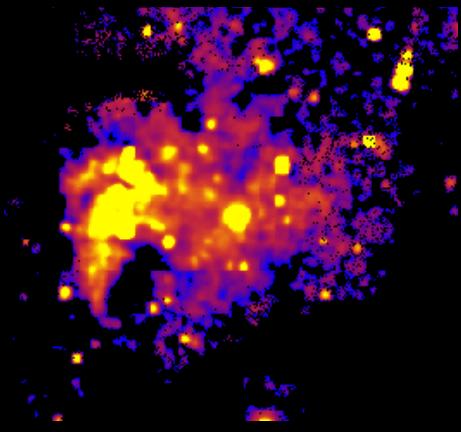
**U4 (Rocket)**



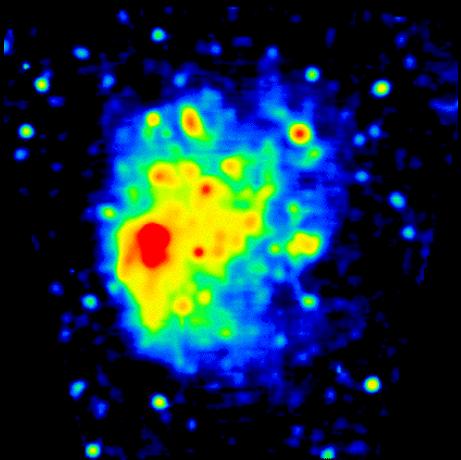
**LMC (AAO)**



**F3 (Herschel)**

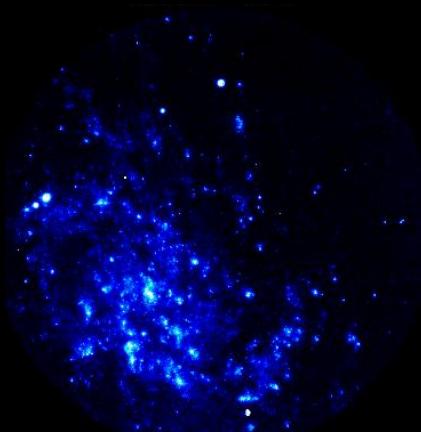


**X1 (ROSAT)**



**R10 (RAIUB)**

# Triangulum



**U1** (UIT)

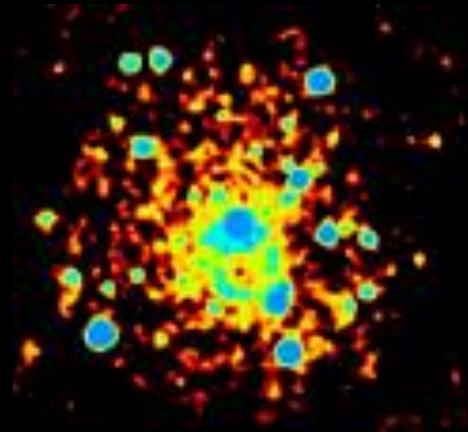


© Malin/IAC/RGO

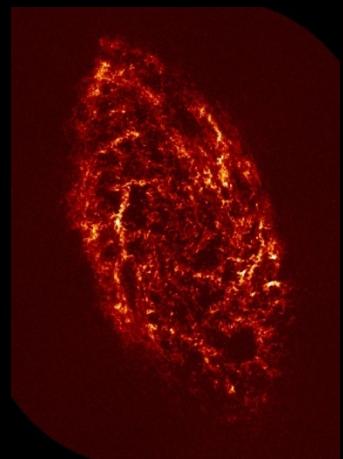
Triangulum (INT)



**M4** (Spitzer)



**X8** (ROSAT)



**R2** (NRAO)

# Orión



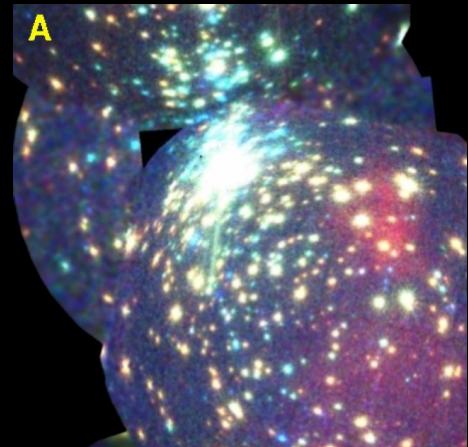
**N1** (VISTA)



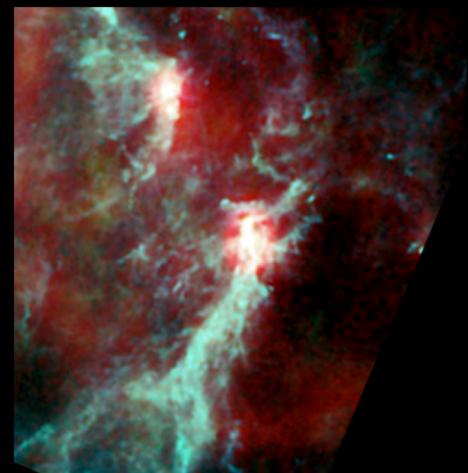
Orion (Hubble)



**M3** (Spitzer)

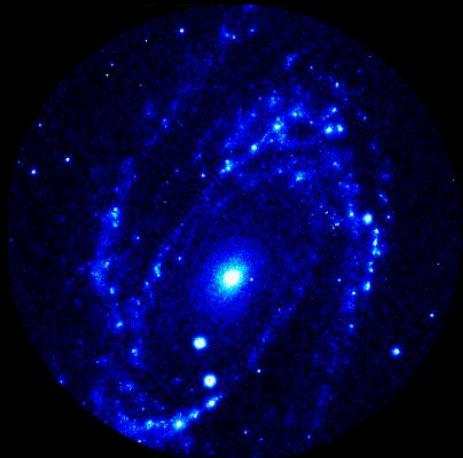


**X2** (XMM)



**F9** (Planck)

# M81



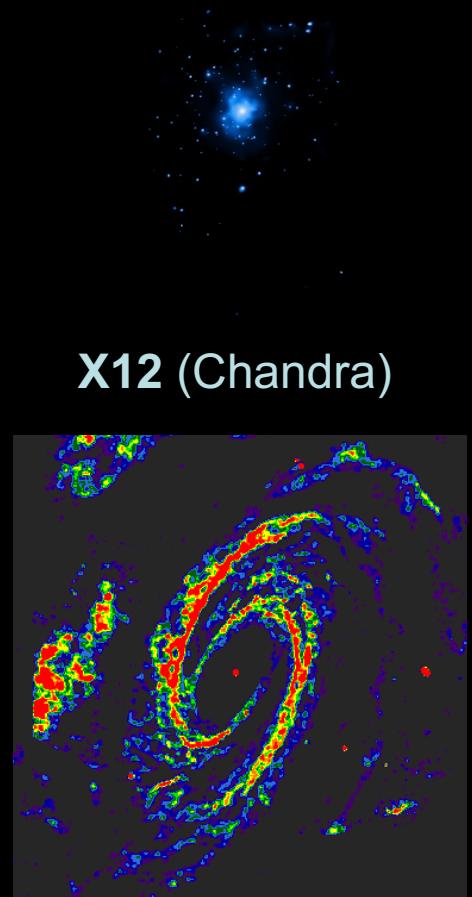
**U2** (UIT)



**F5** (Herschel)

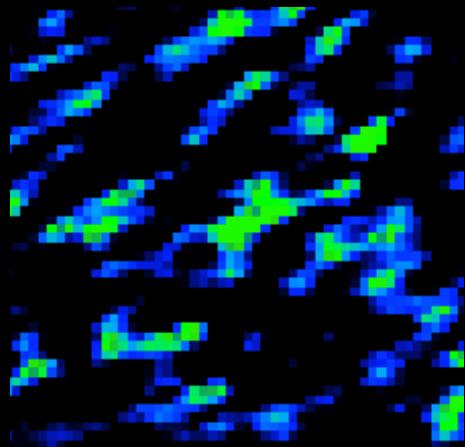


M81 (Gendler)



**X12** (Chandra)  
**R1** (VLA)

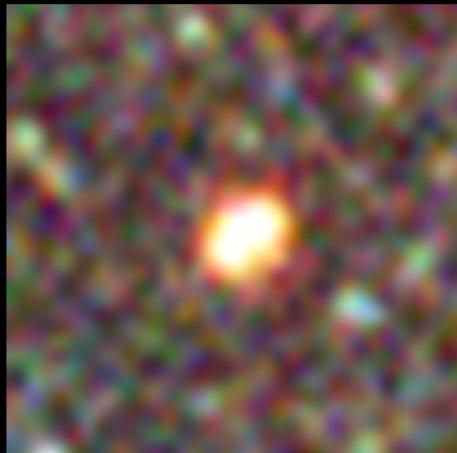
# M87



**M5** (IRAS)



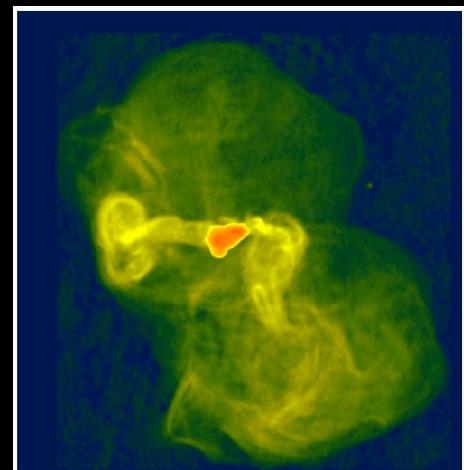
**M87** (AAT)



**F4** (Herschel)



**X9** (Chandra)

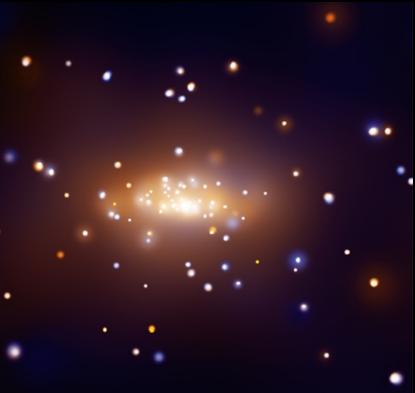


**R4** (NRAO)

# Sombrero



N2 (2MASS)

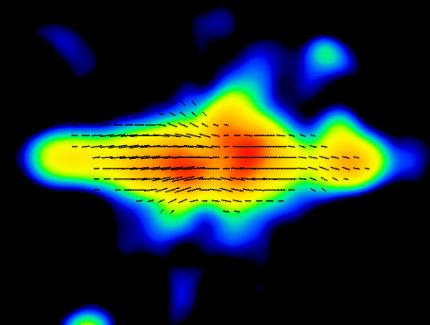


X10 (Chandra)



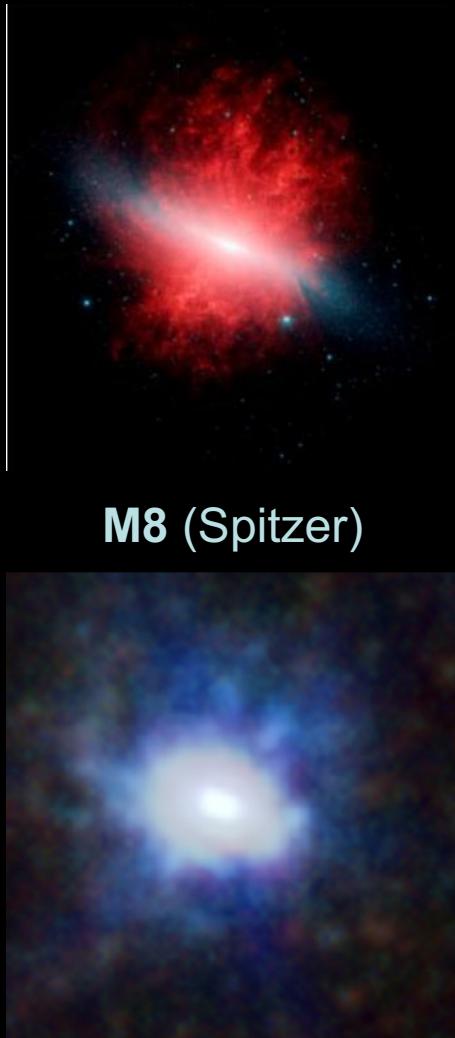
M1 (Spitzer)

Sombrero (Hubble)



R3 (VLA)

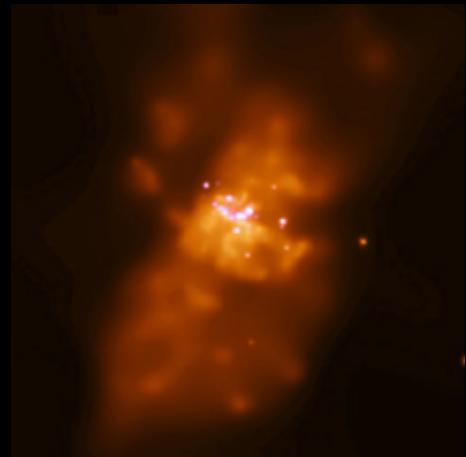
# M82



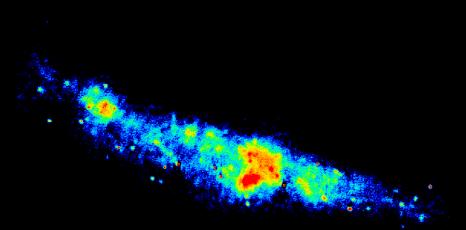
**M8** (Spitzer)



M82 (Gendler)



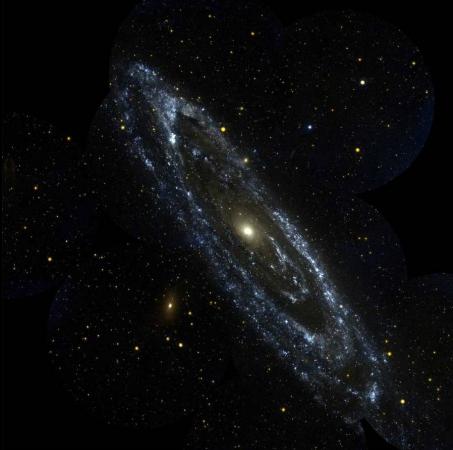
**X11** (Chandra)



**F6** (Herschel)

**R5** (VLA/Merlin)

# Andrómeda



**U3** (GALEX)



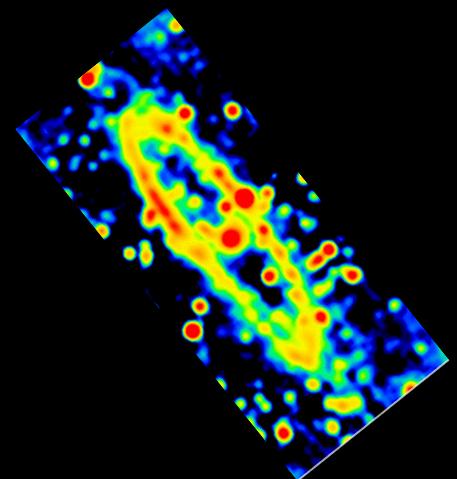
Andrómeda (Gendler)



**F10** (Herschel)

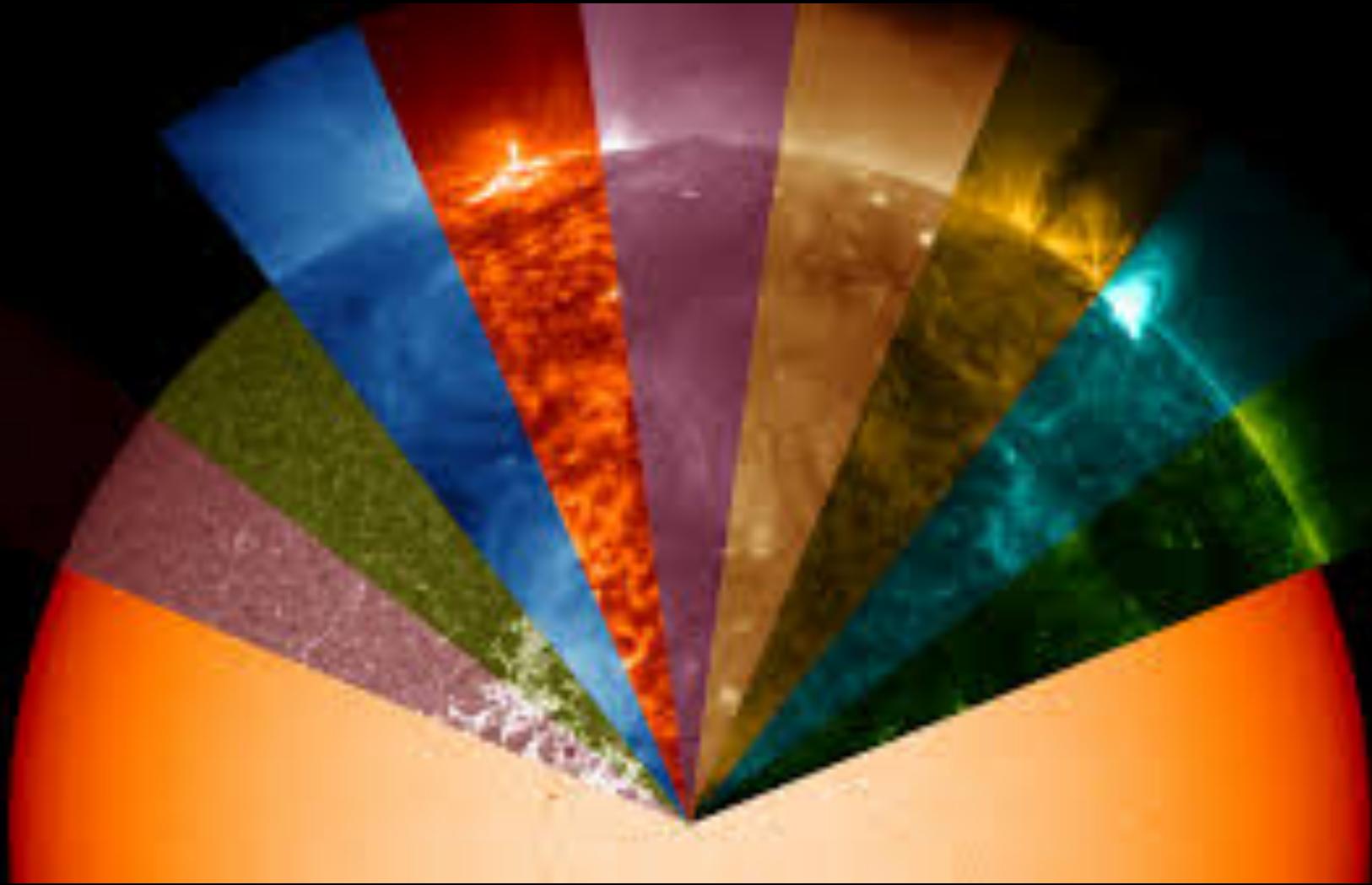


**X7** (XMM-Newton)



**R6** (Effelsberg)

# Rebanadas



<http://www.chromoscope.net>